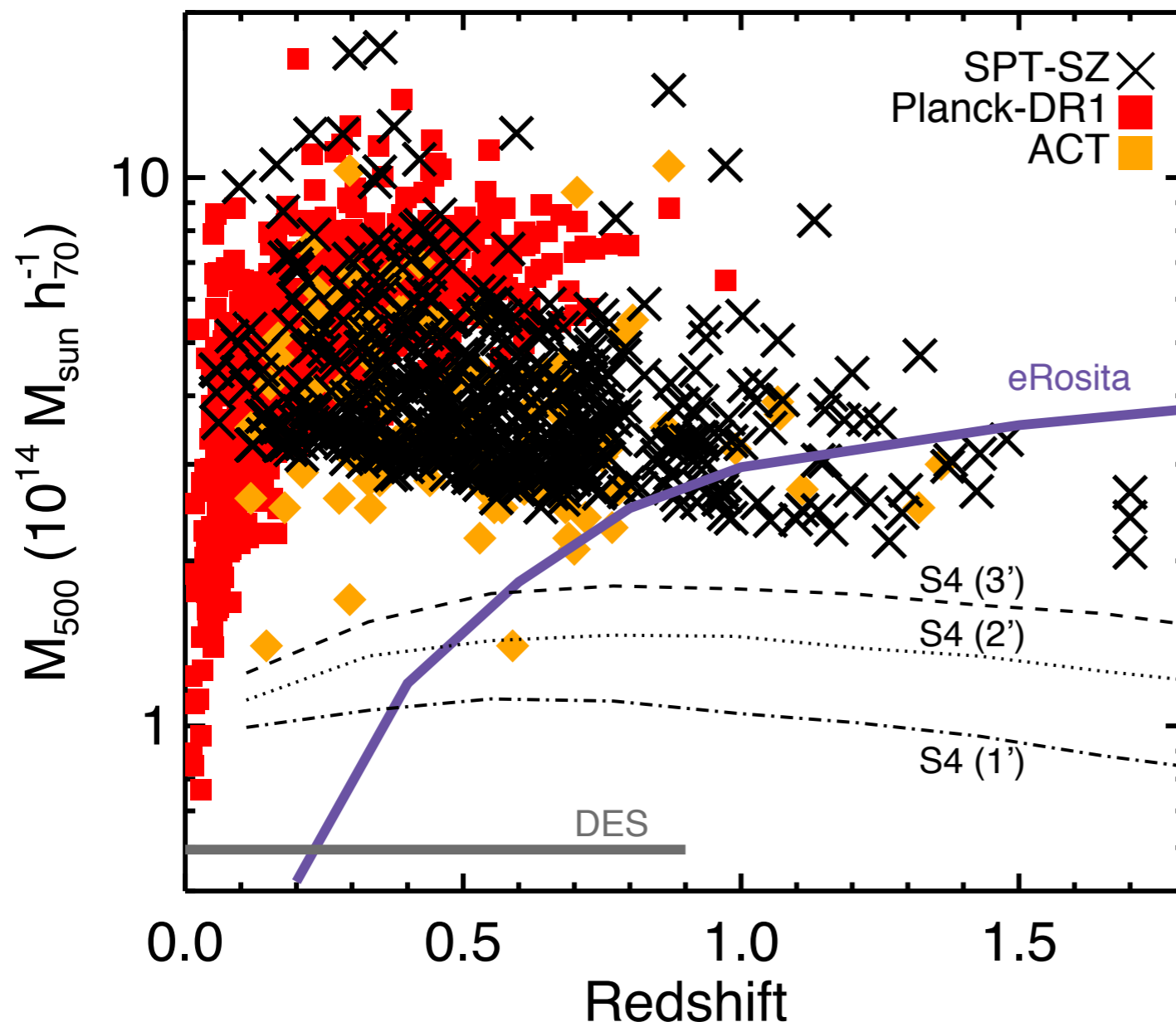


Cross-Correlation Science from SPTxLSST

LSST x CMB-S4	CMB Lensing Convergence (κ) Map	SZ Thermal	CMB Polarization	Small-scale CMB Lensing
Galaxy Map	Measure galaxy bias; the relationship between galaxy stellar and halo mass	Joint cosmology from SZ+Optical Cluster Survey	Gravitational potential from LSST galaxies convolved with E-modes to estimate their B-mode contribution	CMB cluster lensing of LSST galaxies
Shear Map	Constrain systematics in galaxy shape measurements (e.g., linear and magnification bias)	Probes baryonic pressure as a function of distance from halos	Gravitational potential from LSST shear convolved with E-modes to estimate their B-mode contribution	Measure CMB lensing around LSST shear peaks, tomographic measurement of large-scale structure
Tangential Shear [galaxy-galaxy or cluster lensing]	Measure tangential shear of LSST galaxies around SZ-selected lensing peaks	SZ cluster survey cosmological constraints using LSST WL mass calibration		

CMB-S4 Cluster Survey



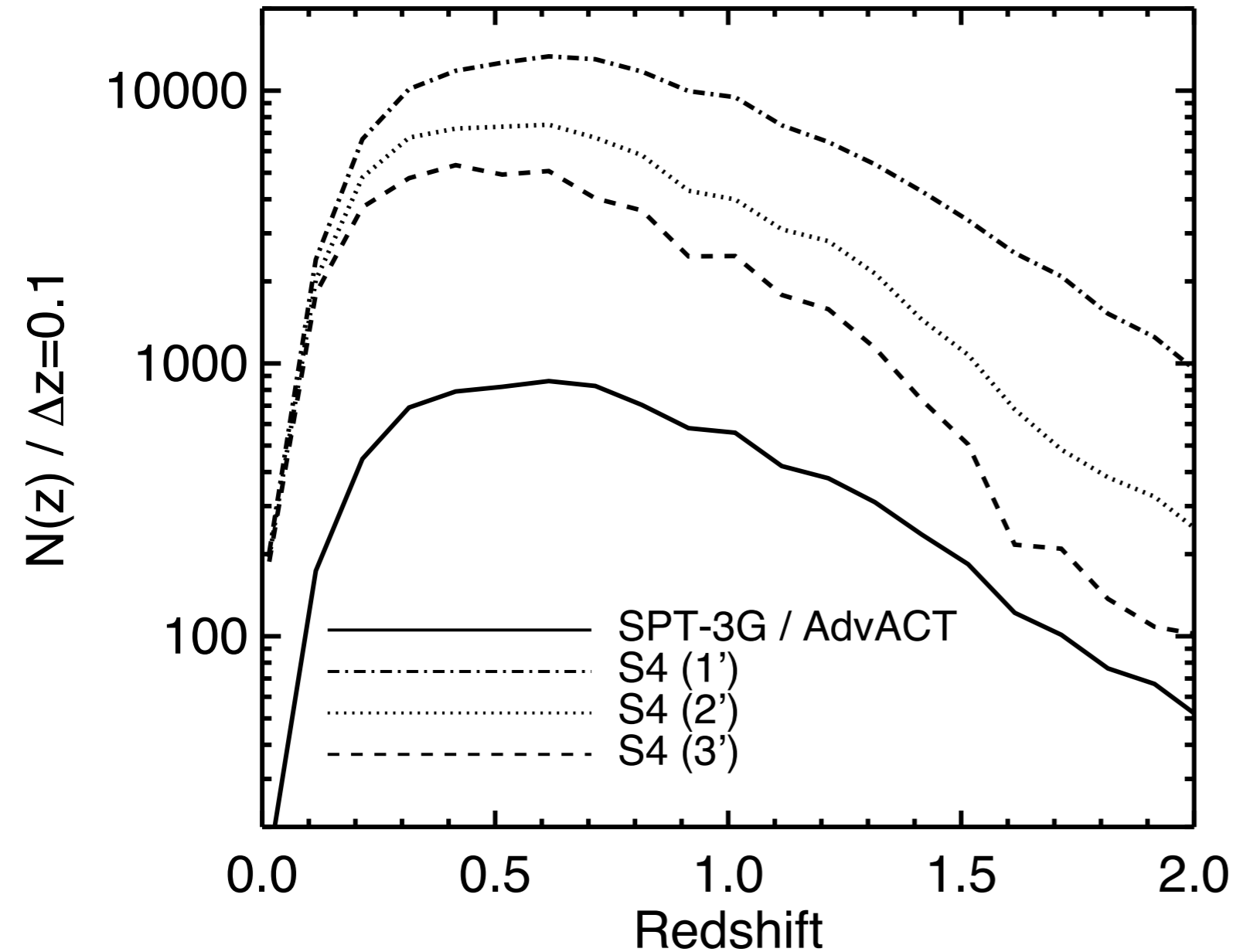
General Parameters of Cluster Survey:

- Estimated 50% mass-completeness of S4 survey
- Assumed $f_{\text{sky}}=0.5$ and 95, 150, 220 GHz bands at depths of 2, 1, 2 $\mu\text{K-arcmin}$, respectively
- Used old Shaw+08 sims, re-scaled SZ amplitude to match SPT-SZ cluster counts
- ***Mass limit for CMB-S4 survey will be ~2-3 times less than current SZ surveys***

Action Items:

- Use new sims (e.g., Battaglia)
- Consider more variations on S4 survey parameters, bands, depths, etc.

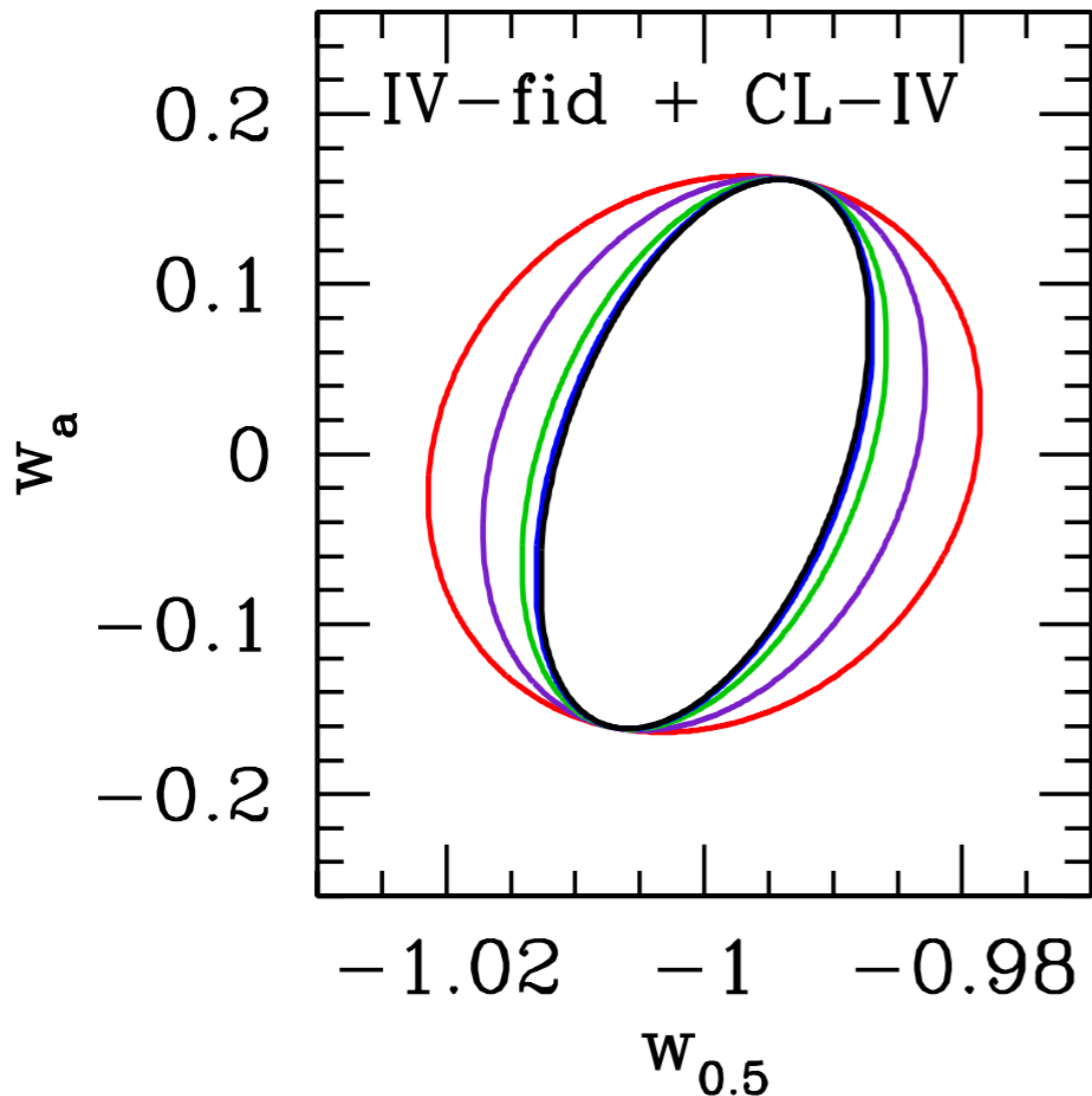
CMB-S4 Cluster Survey



General Parameters of Cluster Survey:

- CMB-S4 will find on order $140e3$, $70e3$, $45e3$ clusters for a 1, 2, 3 arcmin beam, respectively
- From cluster abundance, expect a DETF FOM $\sim 140, 70, 45$, respectively.

CMB-S4 Cluster DE Forecasts

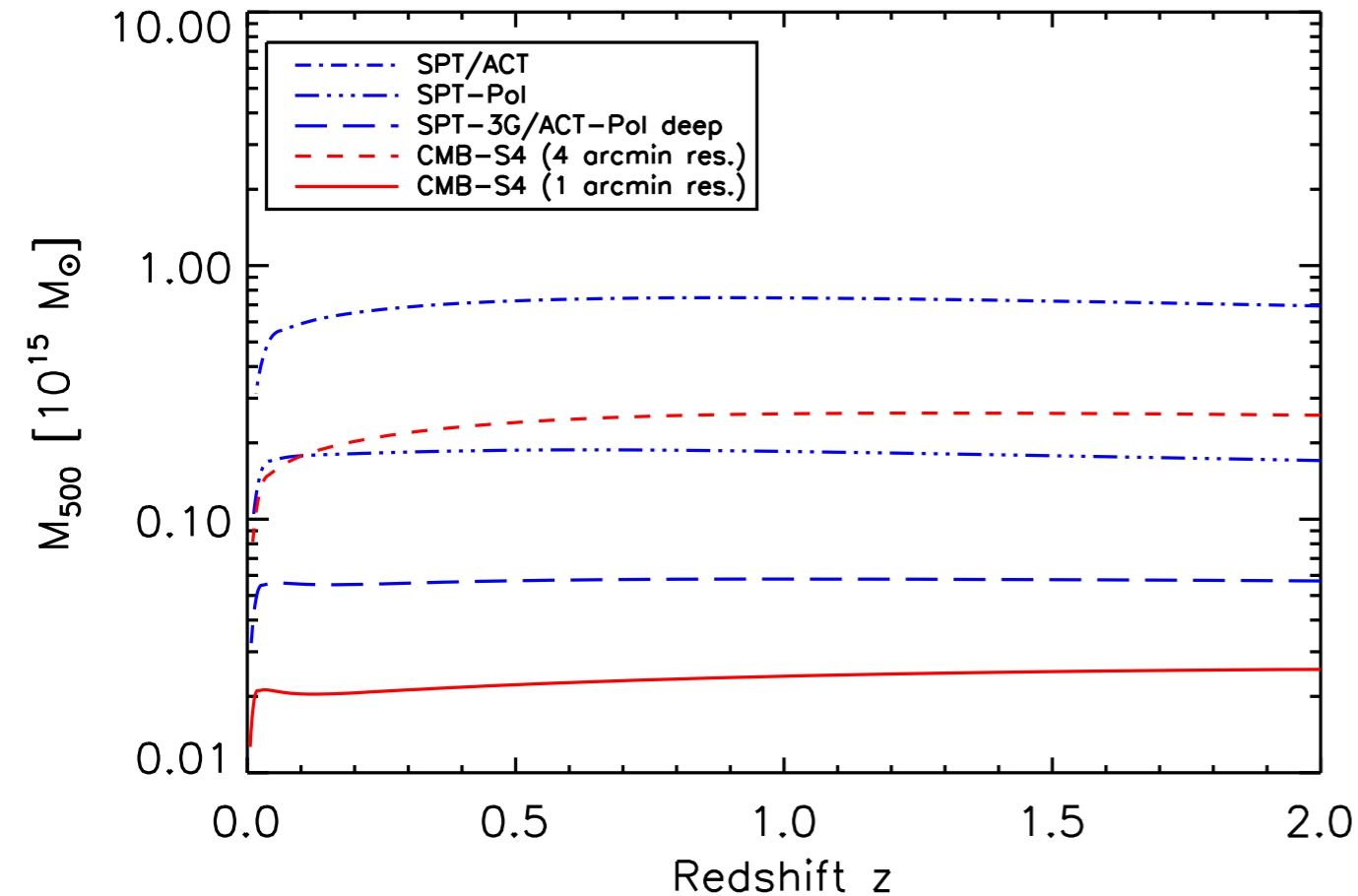


- Weinberg+(2013) found S4-cluster survey could improve DE FOM by factor of ~ 2 , over baseline S4 constraints from BAO, SNe, WL, Planck

Action Items:

- Need DE projections from cluster counts
 - DE improvement from adding CMB-S4 to other S4 surveys (i.e., DESI, LSST, etc.)
 - Some potential candidates to do this work identified, will talk to Eduardo Rozo this week as well
 - DE constraints from joint LSST+S4 cluster survey
 - ***Incorporate into larger CMB-S4 DE projection framework (not clear who is doing this)***

CMB-S4 Cluster Ancillary Science



Case and level of detail in science book is still TBD:

- CMB cluster lensing
- Cluster astrophysics and evolution
 - Unique sample of $\sim 10,000$ massive high- z clusters
 - Formation and gas virialization in the “first” clusters
- Cluster profiles
 - Missing baryons
 - Astrophysical feedback history

Anything else ...?