Fig. 45. Amplitude ratio between total polarized foregrounds and CMB as a function of both multipole moment and frequency, defined by $f(\ell, \nu) = [C_\ell^{\mathrm{fg}}(\nu)/C_\ell^{\mathrm{CMB}}]^{1/2}$, as defined Eq. 22 with parameters derived from 73 % of the sky. The left and right panels show the *EE* and *BB* spectra, and the black and red contours in the latter corresponds to tensor-to-scalar ratios of $r = 0.0$ and 0.05, respectively.